

An ALMA multi-line survey of the interstellar medium of the redshift 7.5 quasar host galaxy J1342+0928

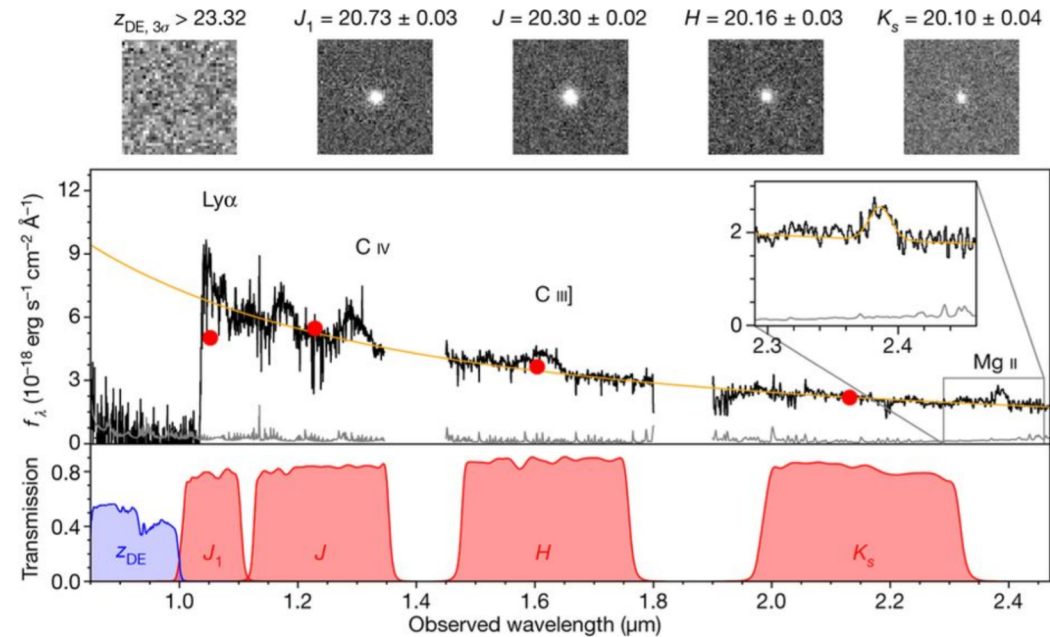
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ULAS J1342+0928

- Quasar of the highest redshift ($z = 7.54$)
- $L_{bol} = 10^{13} L_{\odot}$
- $M_{SMBH} = 8 \times 10^8 M_{\odot}$

ALMA Multi-line Targeting

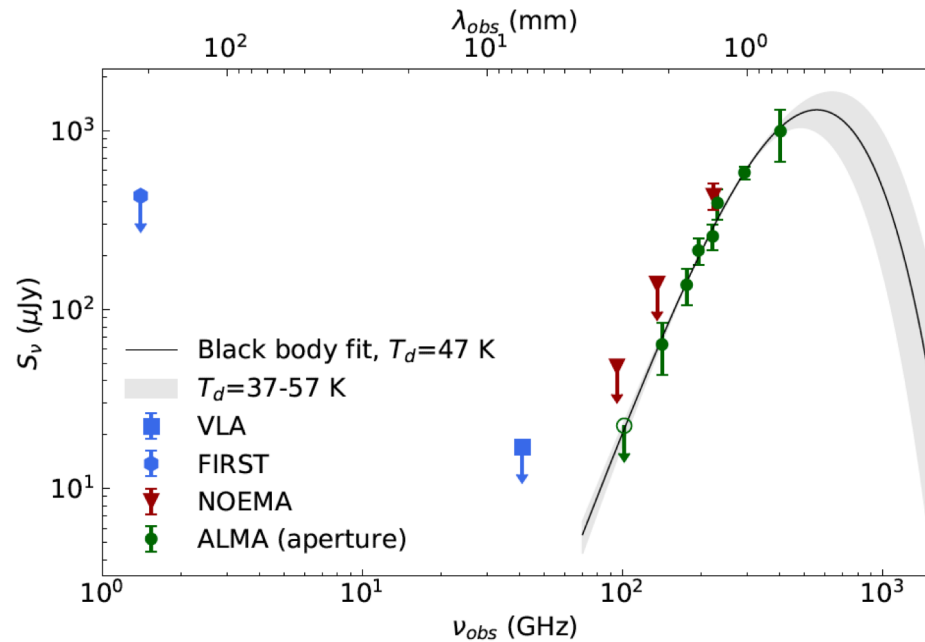
- 24 molecular and fine-structure lines
- eight different frequency located between 93.5 GHz (band 3) and 412 GHz (band 8) with effective bandwidth of 60 GHz



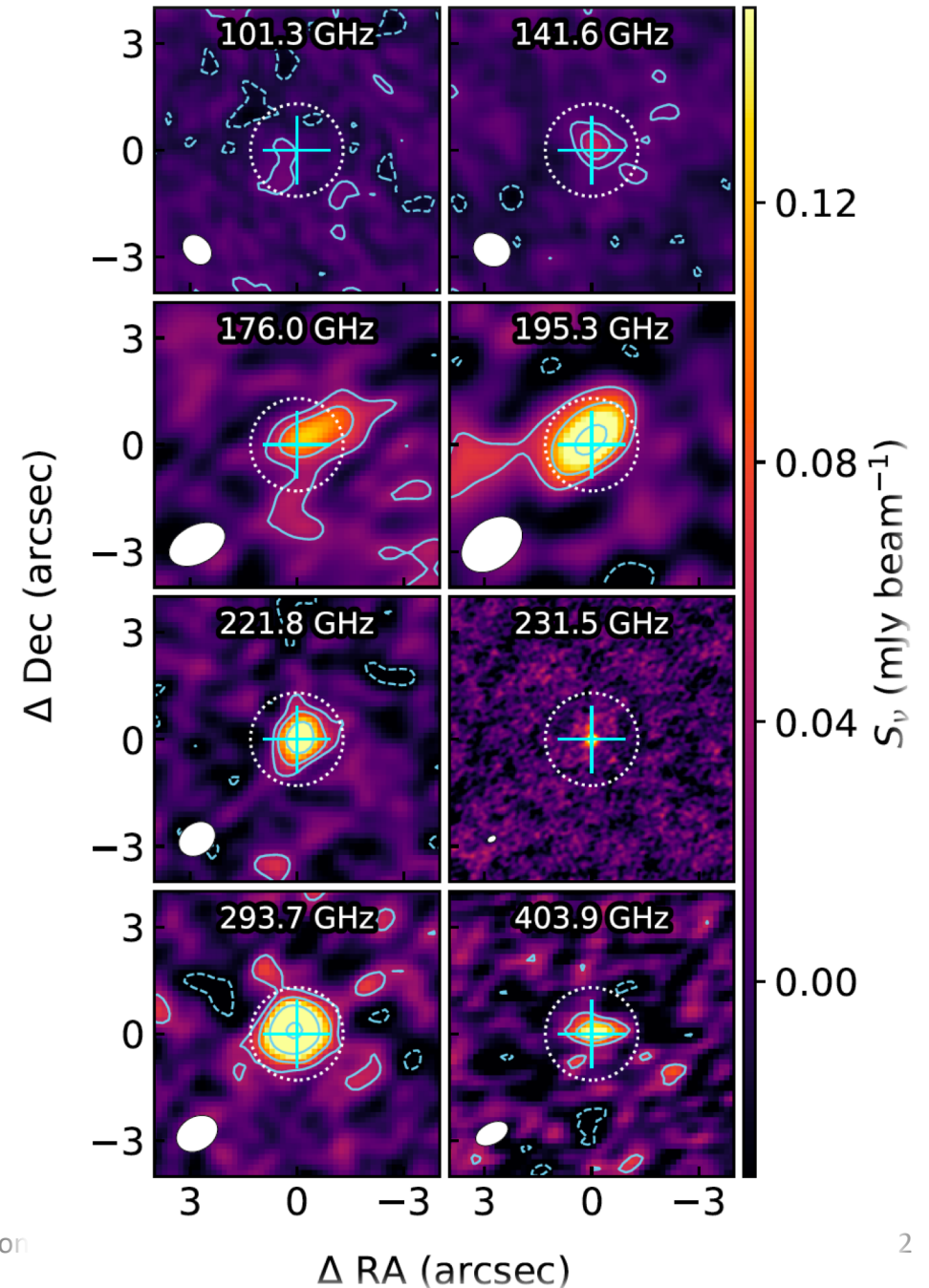
Banados et al.
2018, Nat.

A plethora of information on the ISM condition
in the most quasar host

ULAS J1342+0928: Dust Continuum

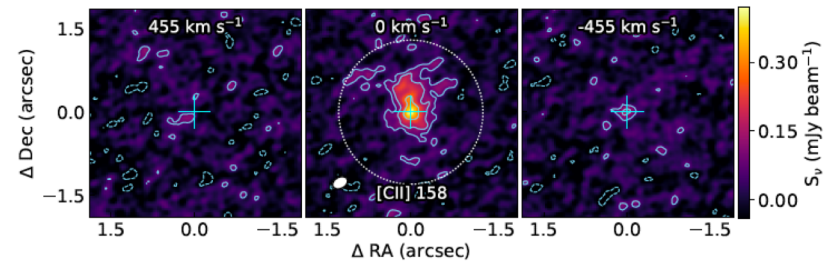


- $M_{\text{dust}} = (0.35 \pm 0.02) \times 10^8 M_\odot$
- $L_{\text{TIR}} = (1.5 \pm 0.3) \times 10^{12} L_\odot$, ULIRG
- $SFR = 150 \pm 30 M_\odot \text{yr}^{-1}$

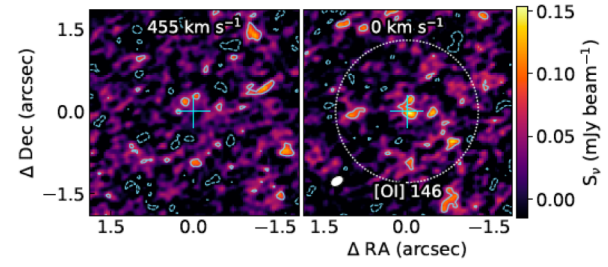


ULAS J1342+0928: Emission Lines

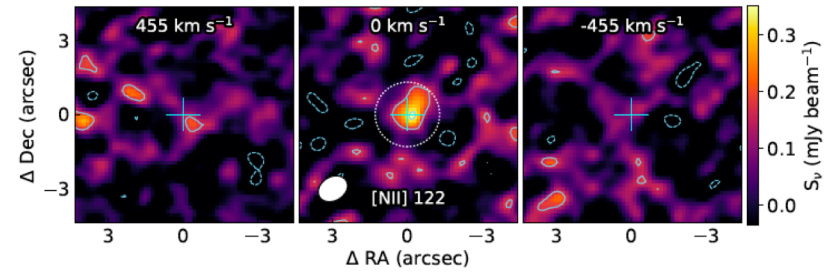
- Low gas density $\lesssim 5 \times 10^4 \text{ cm}^{-3}$ & Strong radiation fields $\gtrsim 10^3 G_0$ ($G_0 = 1.6 \times 10^{-3} \text{ erg cm}^{-2} \text{ s}^{-1}$) for PDR, (from line ratios)
- High electron densities for HII region, $n_e > 180 \text{ cm}^{-3}$, $\sim 16\%$ of hydrogen in ionized form (from N++ lines)
- Low gas-to-dust ratio < 100 (from CO and SLED)
- **Highly enriched**: $Z_{\text{gas}} = 1.3^{+0.3}_{-0.1} Z_{\odot}$ (680 Myr after the BB) (from [OIII] 88/[NII] 122)



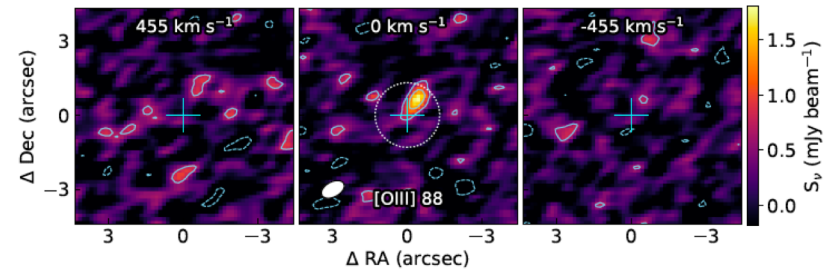
[CII] 158
Brightest,
Meger?



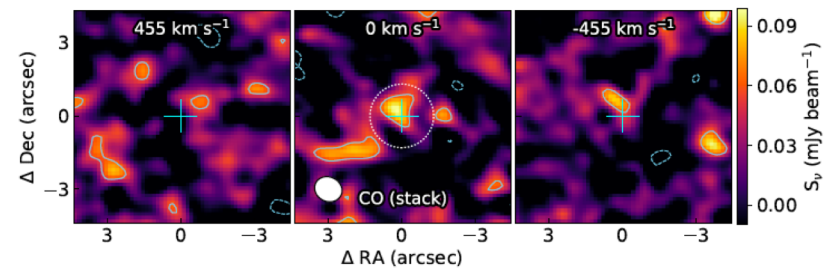
[OI] 146



[NII] 122



[OIII] 188



CO stack